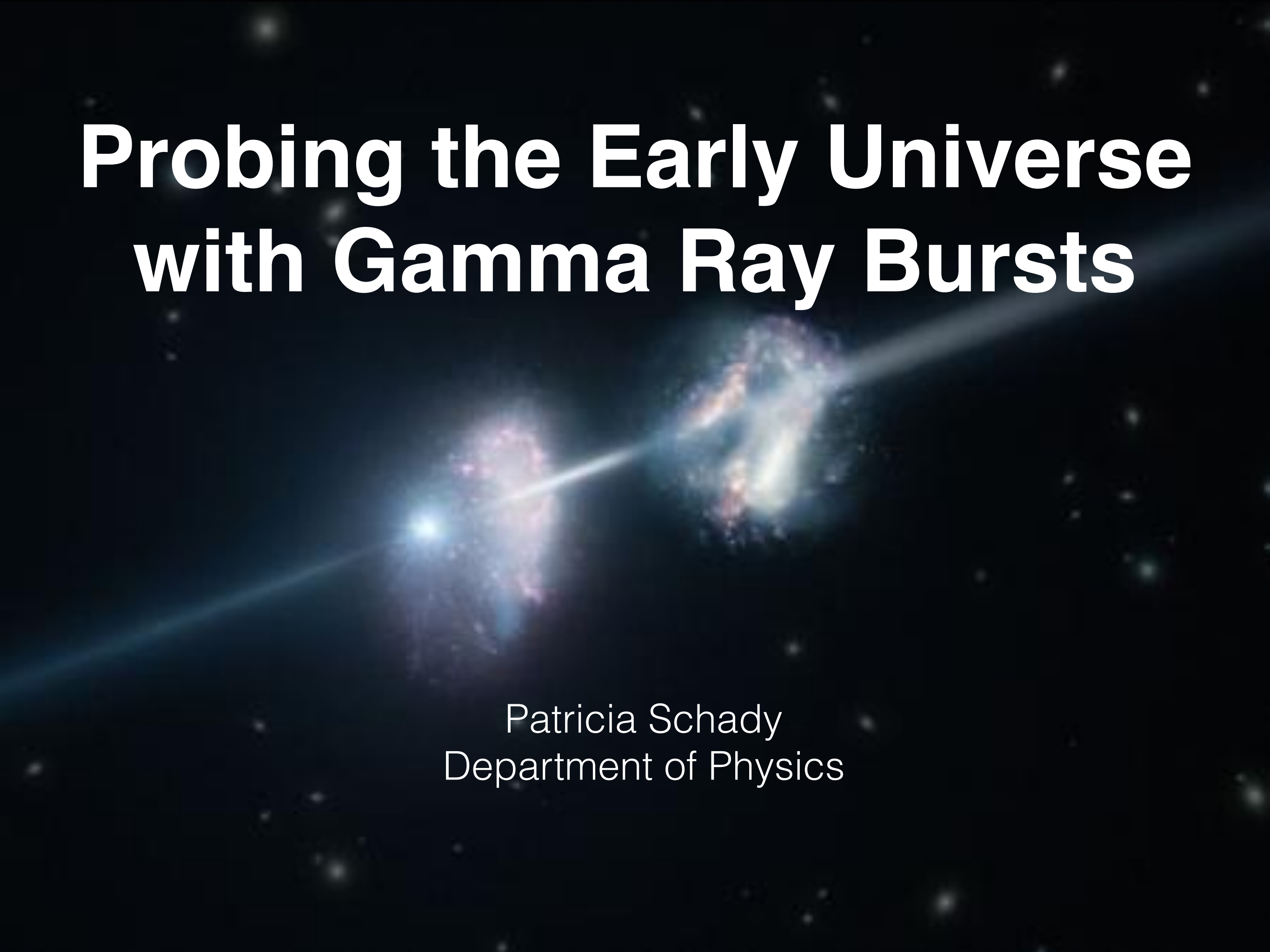
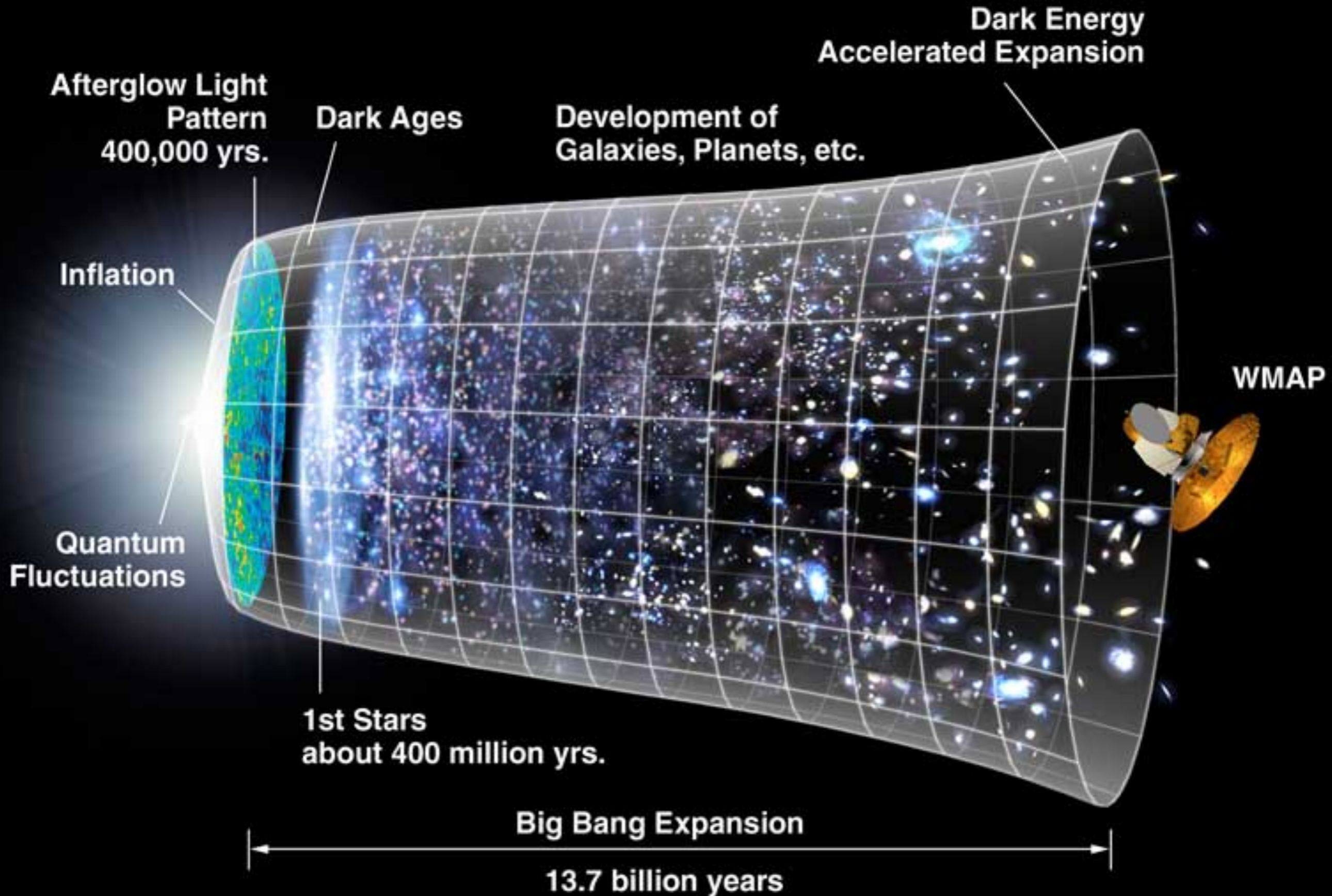


Probing the Early Universe with Gamma Ray Bursts

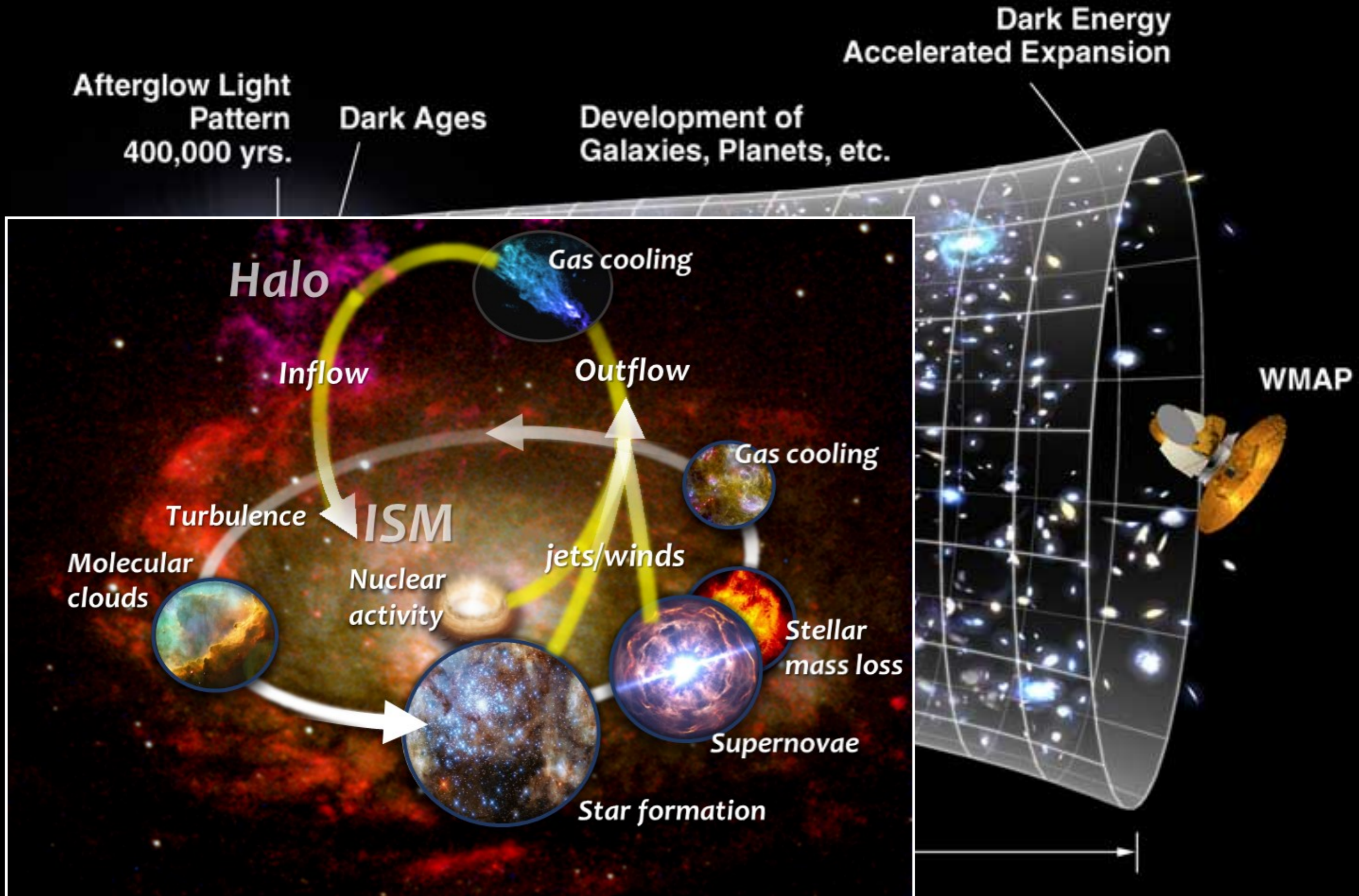


Patricia Schady
Department of Physics

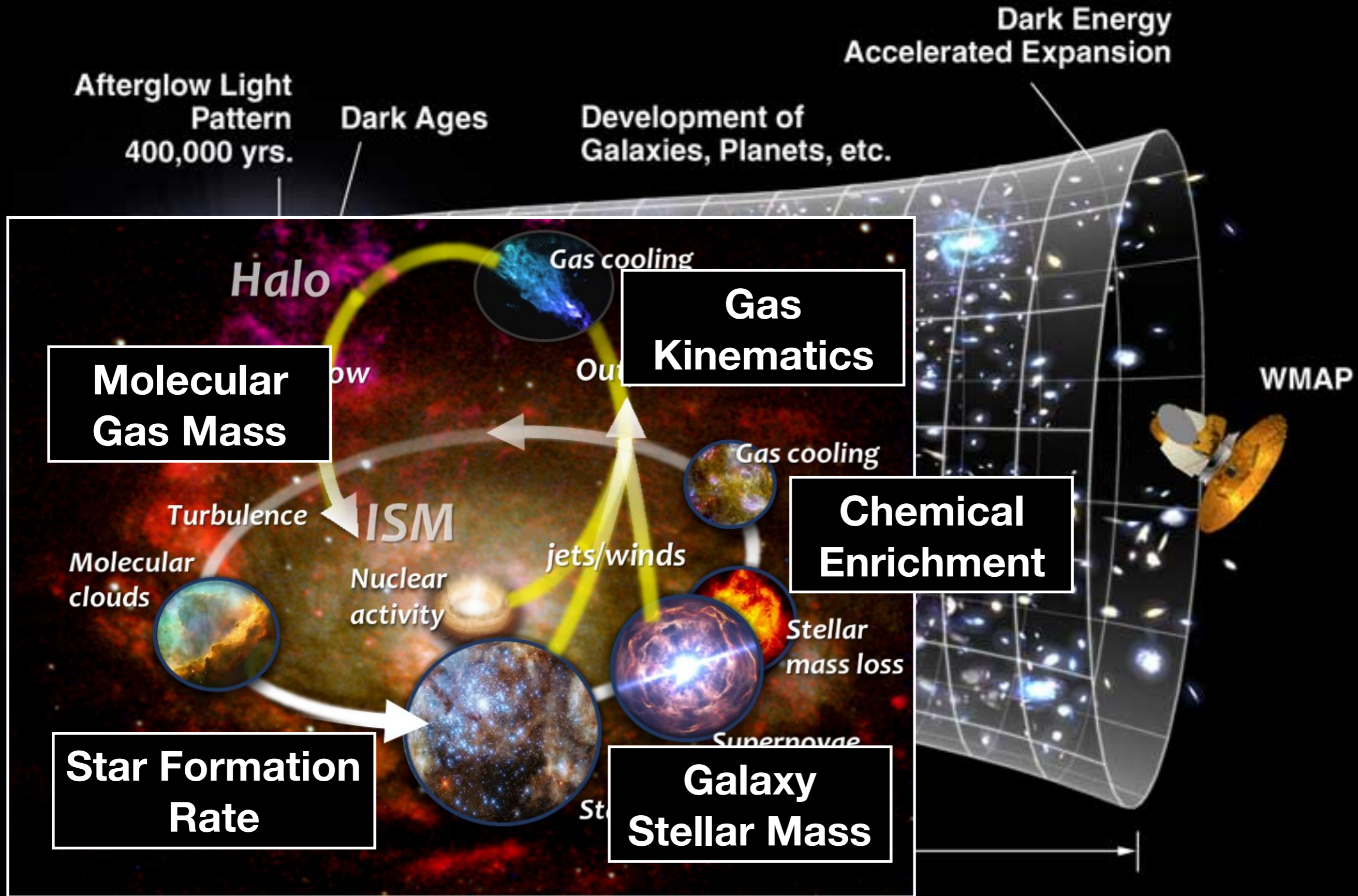
History of the Universe



The galactic baryon cycle

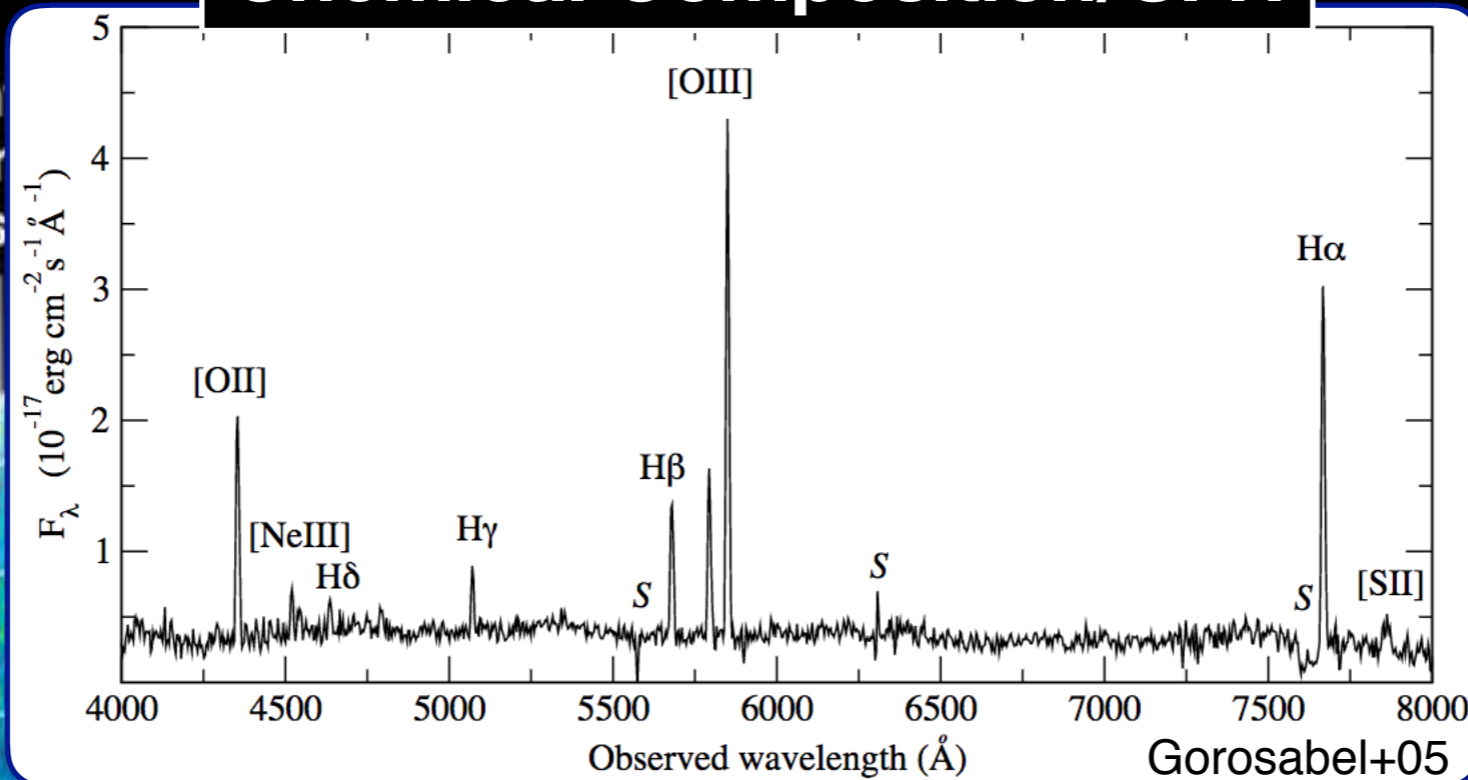


The galactic baryon cycle



Observations of nearby galaxies

Chemical Composition/SFR



Stellar Mass/SFR



13.7 billion years

Dark Energy expansion

Afterglow Light Pattern
400,000 yrs

Inflation

Quantum Fluctuations

1st Stars about 400

WMAP

Credit: Robert Gendler

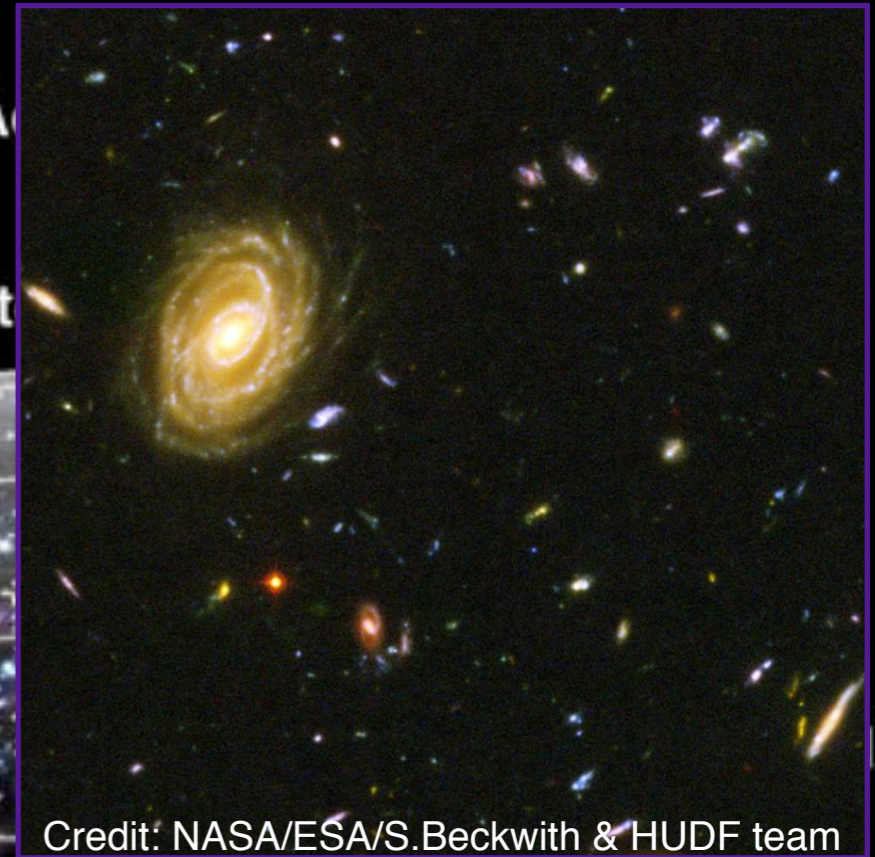
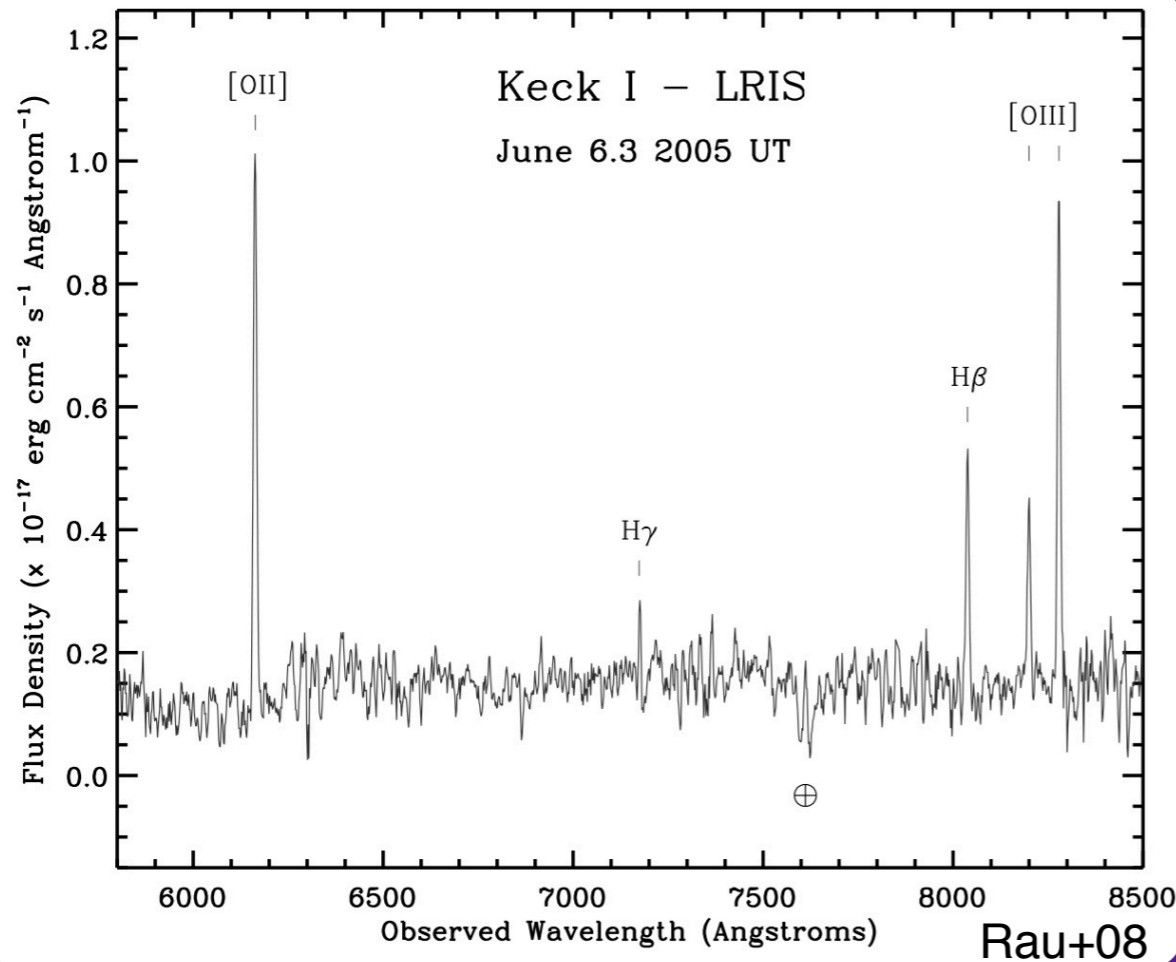
Increasingly distant galaxies

Afterglow Light
Pattern
400,000 yrs.

Dark Ages

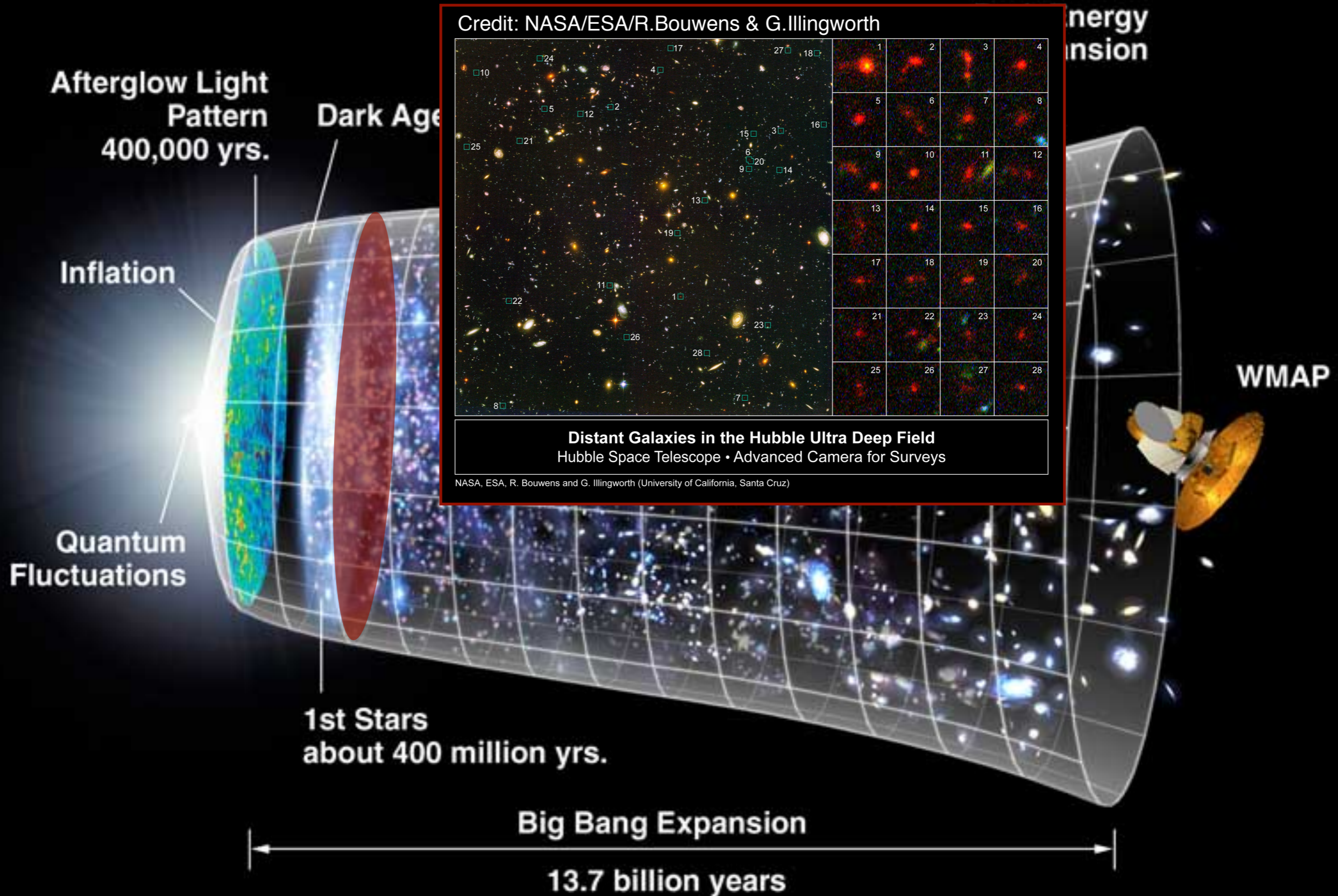
Development of
Galaxies, Planets, et

Inflation



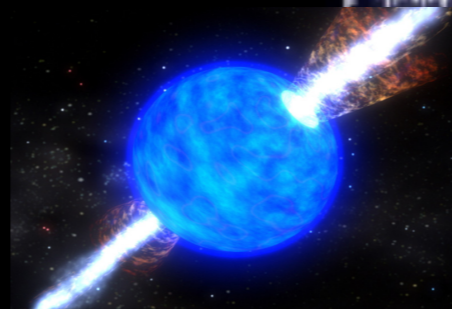
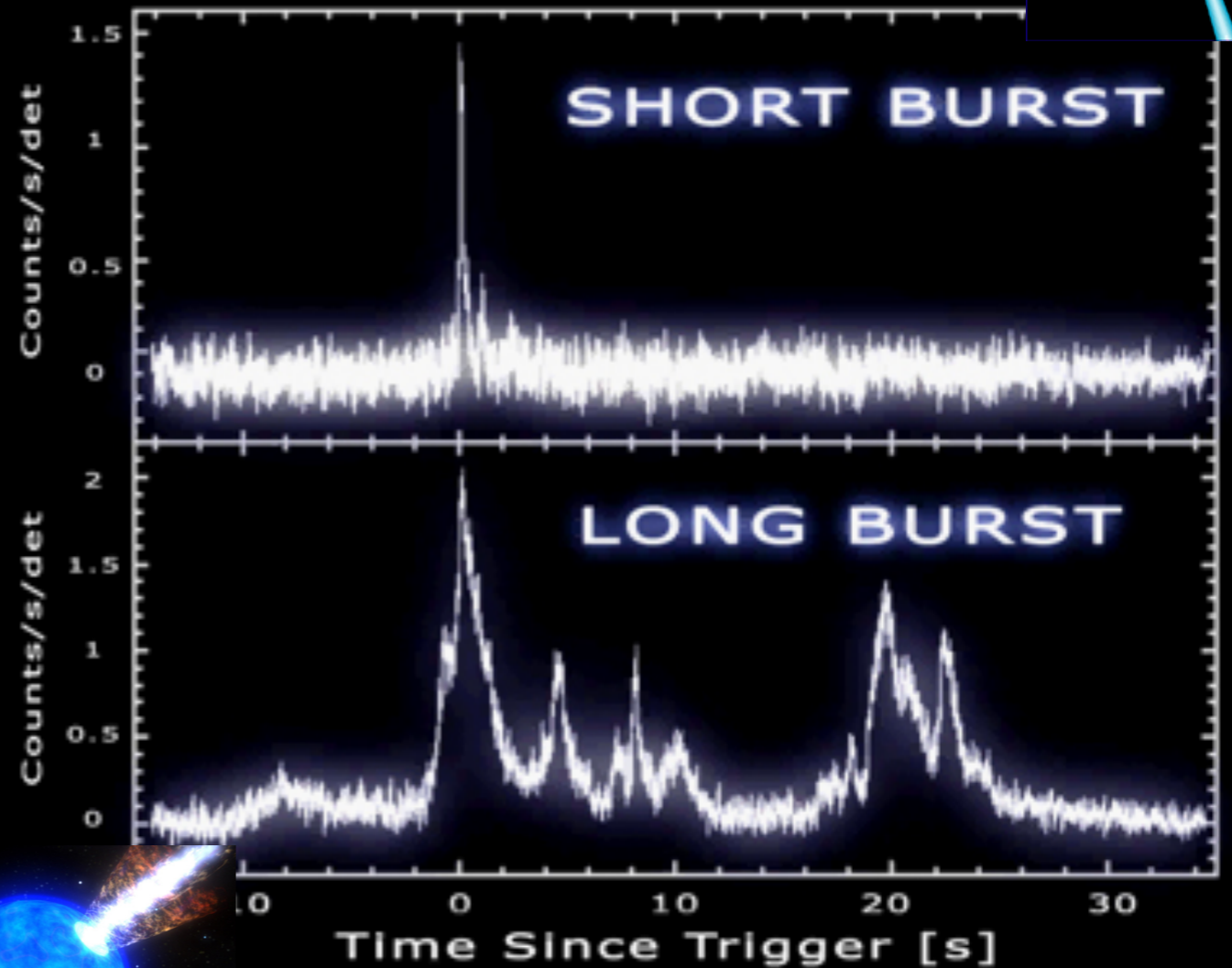
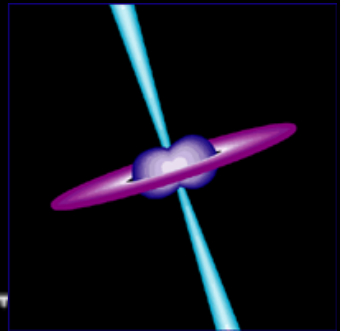
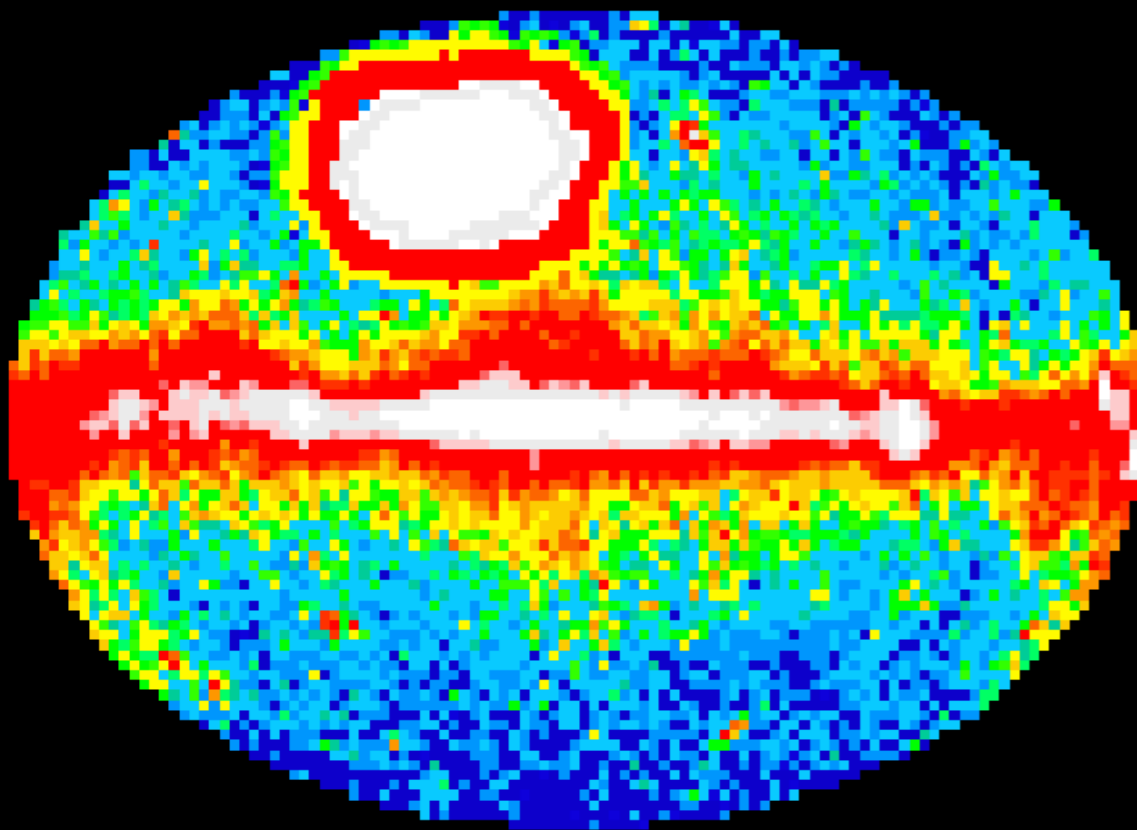
Credit: Robert Gendler

Very distant galaxies



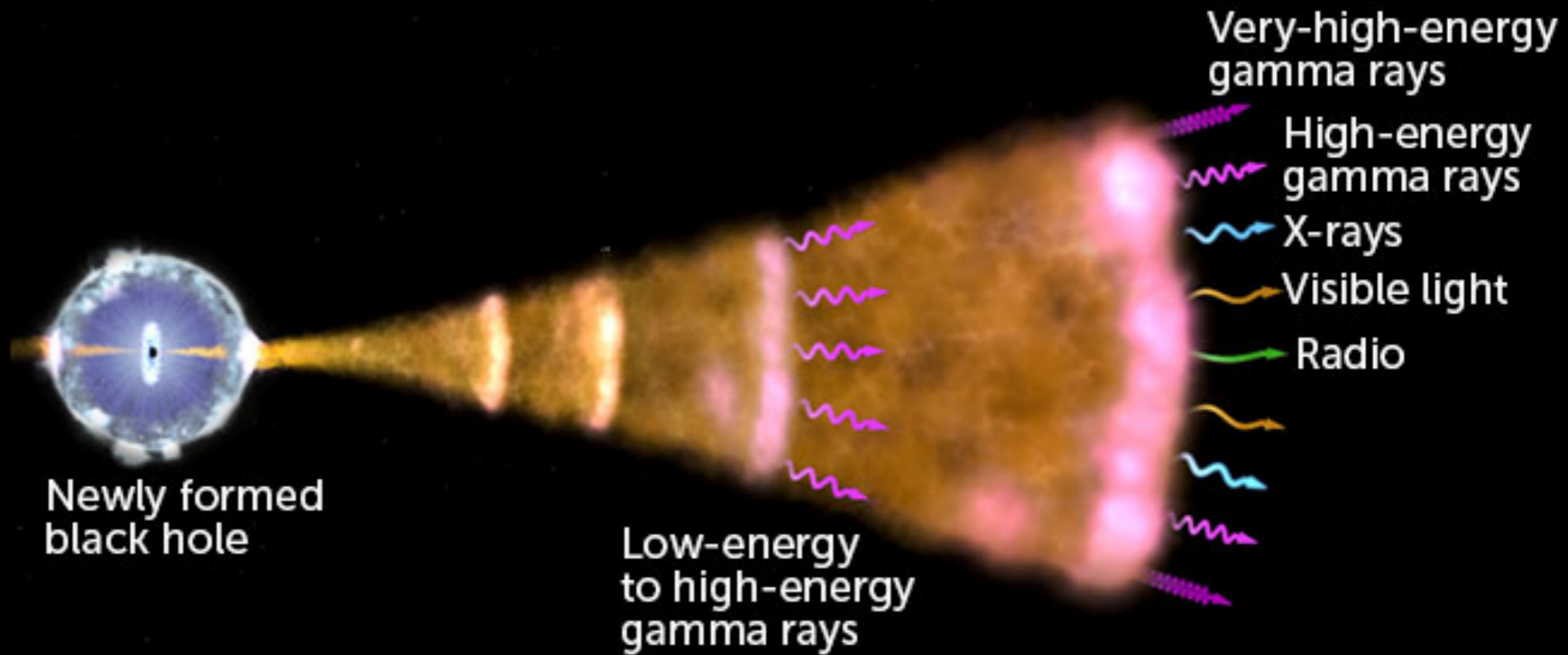
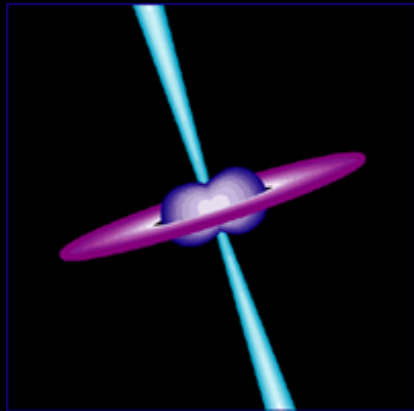
Using GRBs as a probe of cosmic chemical build-up

Gamma Ray Bursts (GRBs)



GRB emission mechanism

Compact star mergers



Newly formed black hole

Low-energy to high-energy gamma rays

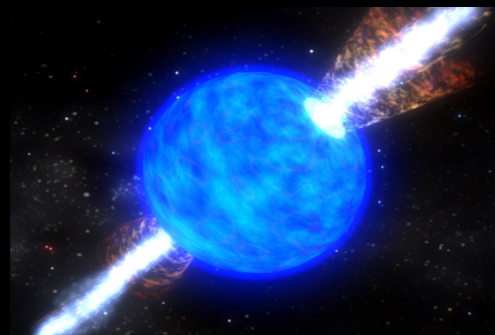
Very-high-energy gamma rays

High-energy gamma rays

X-rays

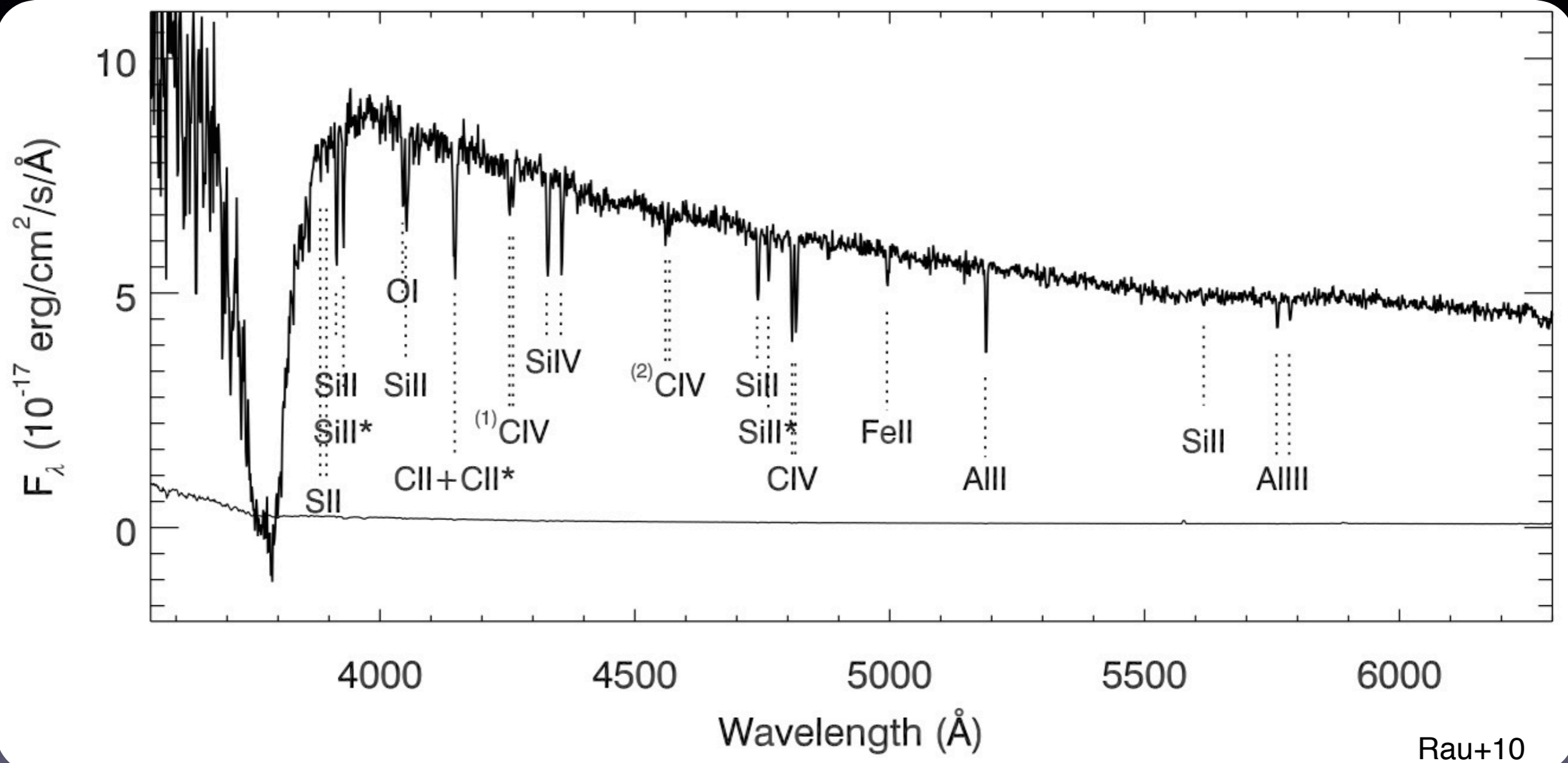
Visible light

Radio



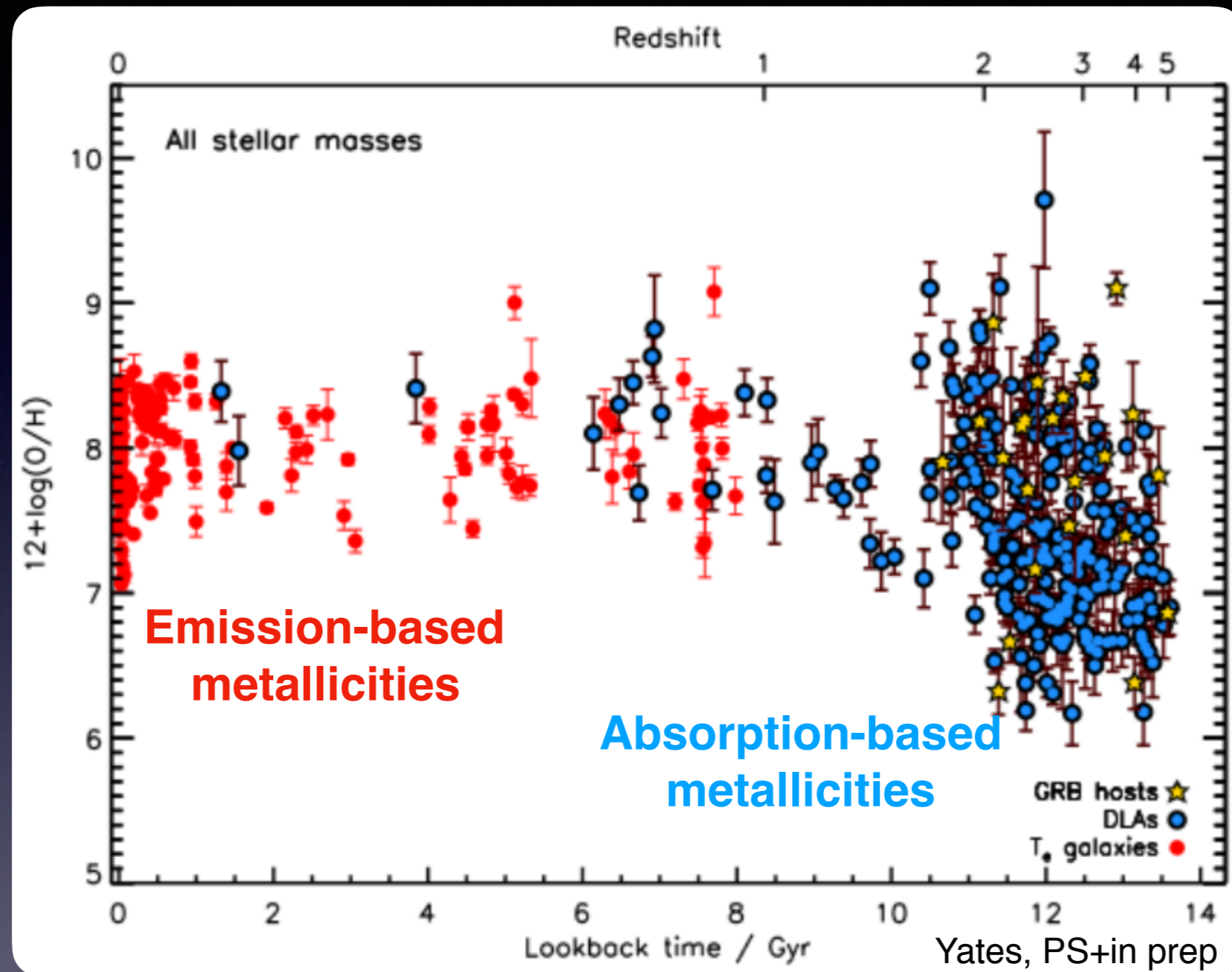
Core collapse of massive star

Metallicities in absorption



Absorption-derived metallicities are largely model-independent and sensitive down to very low metallicity values

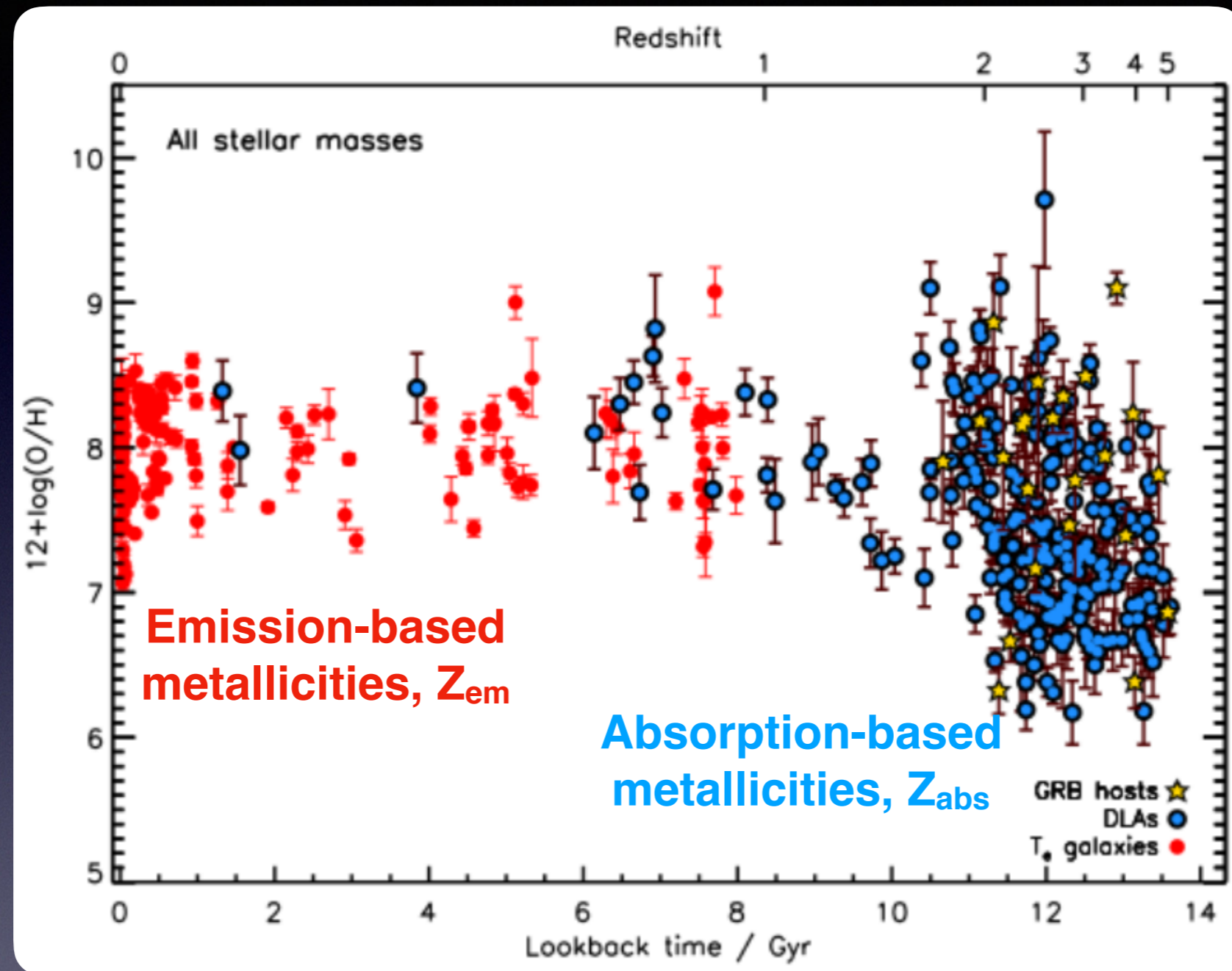
Cosmic chemical evolution



Want to....

- measure host galaxy stellar mass and SFR
- cross calibrate absorption to emission-line derived metallicities

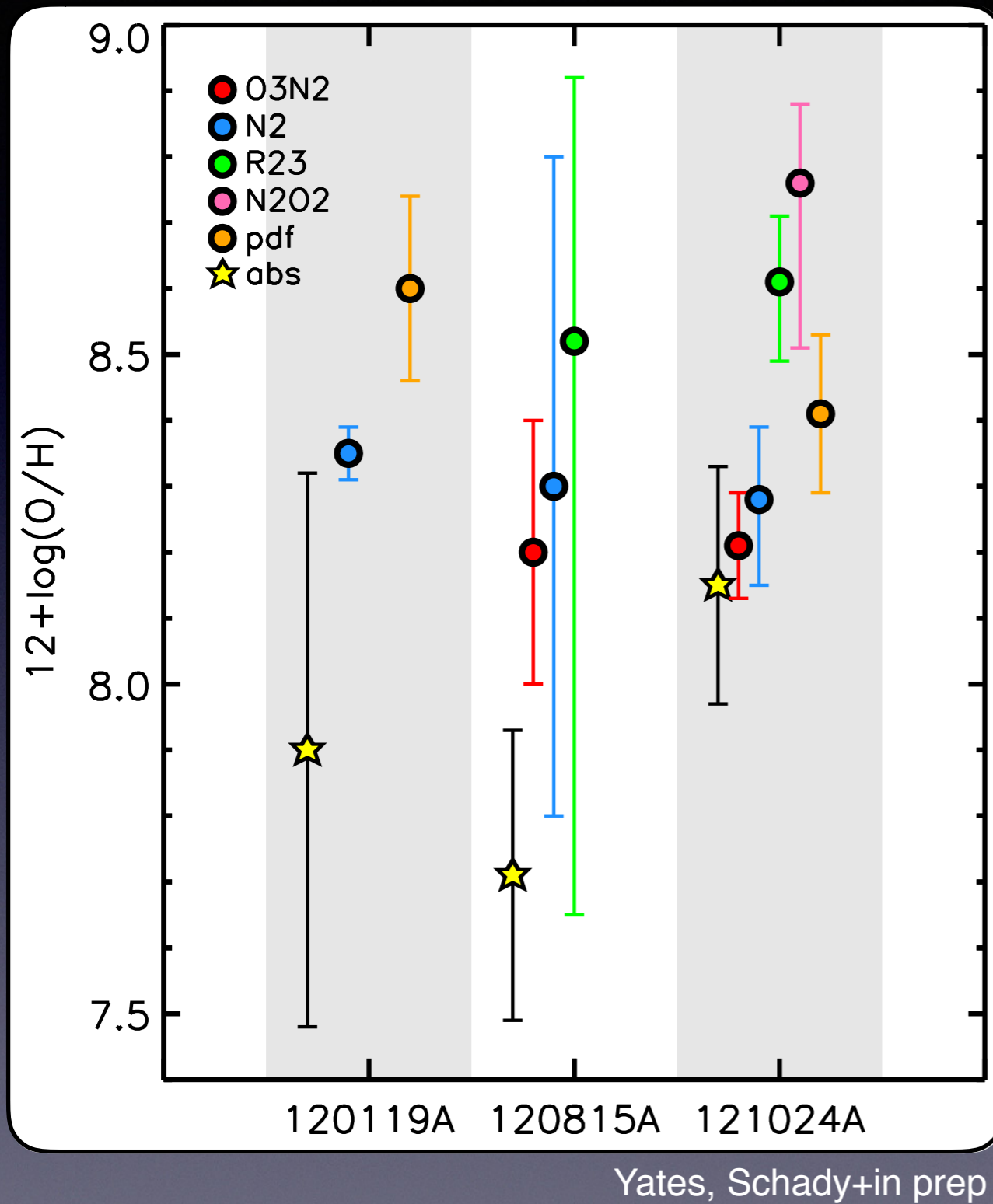
Cross-calibrate metallicity diagnostics



Z_{em} : More sensitive for more massive, chemically evolved galaxies

Z_{abs} : More sensitive for low-mass, chemically unenriched galaxies

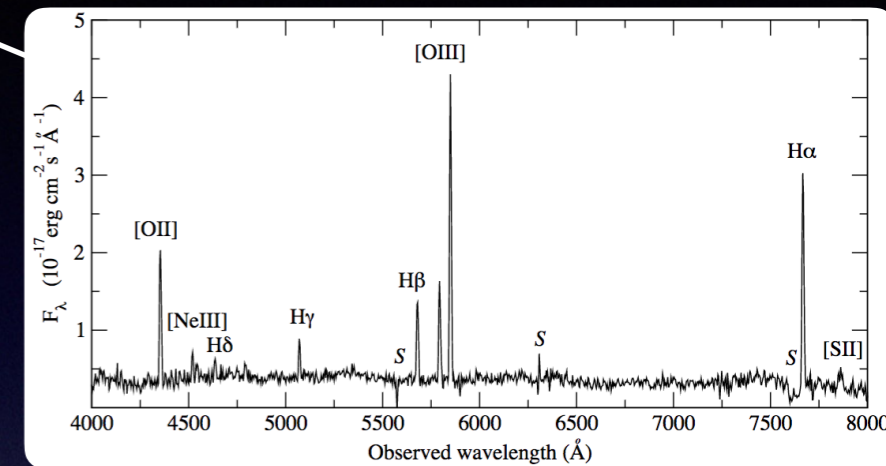
Cross-calibrate metallicity diagnostics



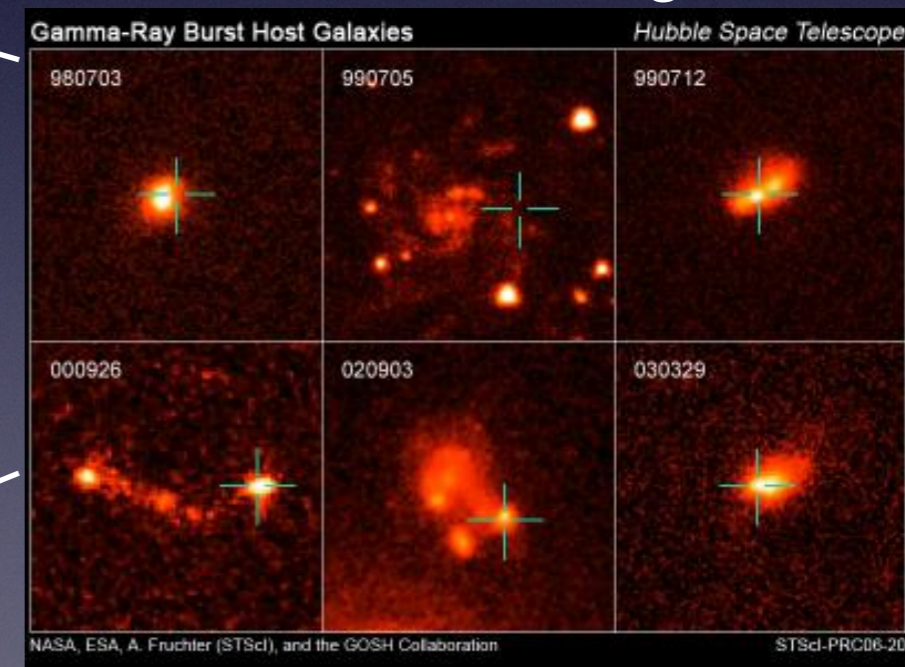
Emission-based
metallicities

Absorption-based
metallicities

Highly model dependent



Local rather than global



With JWST+GRBs can study discrepancy in metallicity diagnostics

Summary



- Long GRBs are excellent and unique probes of distant, star forming galaxies
- The imprint left on their afterglow spectra from intervening material provides a unique view of the chemical composition and conditions of the ISM in their host galaxies
- To optimise their use as probes of distant galaxies, need to investigate how properties inferred from absorption spectroscopy compare to emission-based analysis
- The future NASA JWST mission (all \$9.66 of it!) promises significant headway in this area or research, and in the cosmic chemical evolution in general
- Future, sensitive infrared and X-ray missions (Theseus, ATHENA) will greatly increase high-z GRB samples